

Athena as an Observatory in the late 2020s

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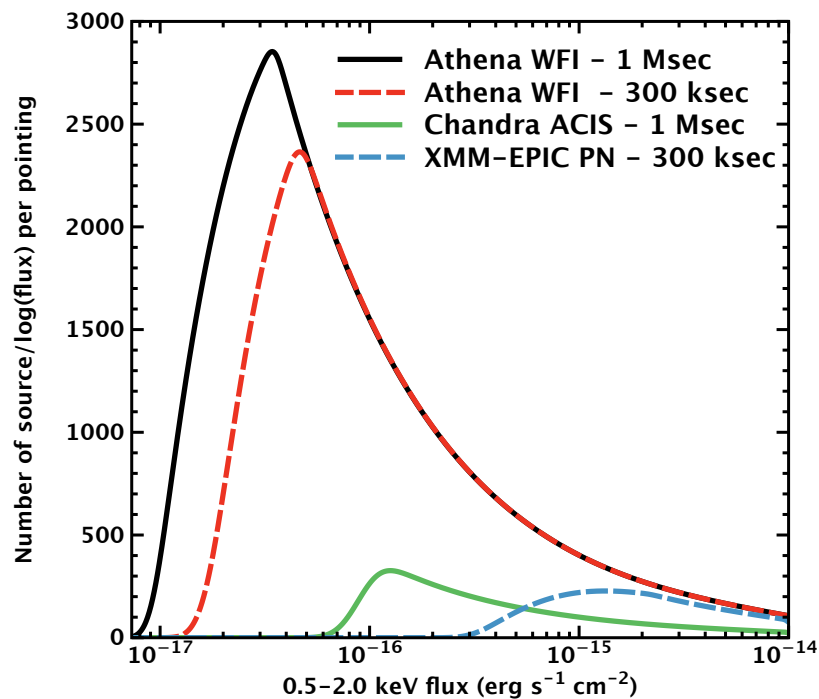
Athena beyond

The Hot and Energetic Universe

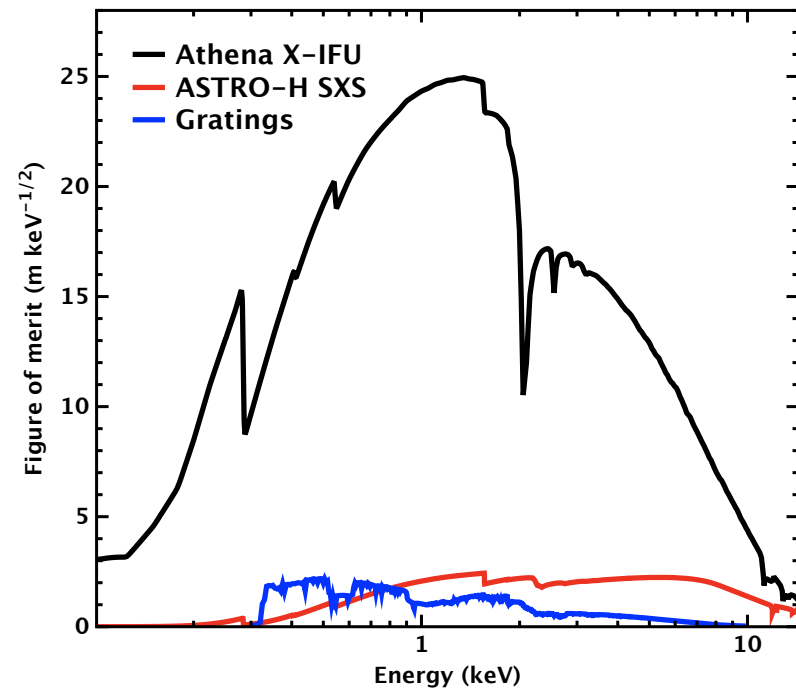
- Athena will address a very challenging and timely set of science objectives on the Hot and Energetic Universe
- But also a large number of **observatory science** capabilities:
 - Solar system planet atmospheres, probing the solar wind
 - Exoplanets: interaction of stellar activity with planets
 - Stellar physics: YSOs, brown dwarfs, massive stars
 - Supernovae & Supernova remnants
 - Stellar end points: neutron stars, magnetars, white dwarfs etc
 - Sgr A*
 - Interstellar medium
 - High-energy transients
- **And a great deal of science that we cannot foresee today**

Athena performance

Survey speed

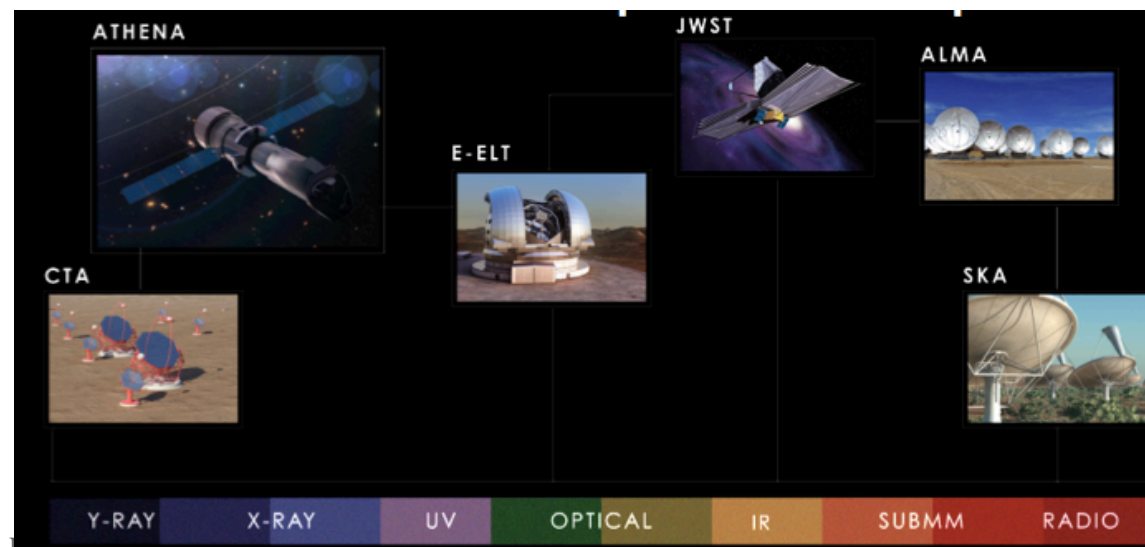


Spectroscopic grasp



The (European) astronomical landscape in the late 2020s

- ASTRONET just released an updated Science Vision and Roadmap
- Major elements:
 - In place: ALMA
 - Approved/in development: JWST, E-ELT, Euclid, Athena
 - In the horizon: CTA, SKA
 - Other: LSST, WFIRST





NINS

ALMA: The global (sub)mm facility



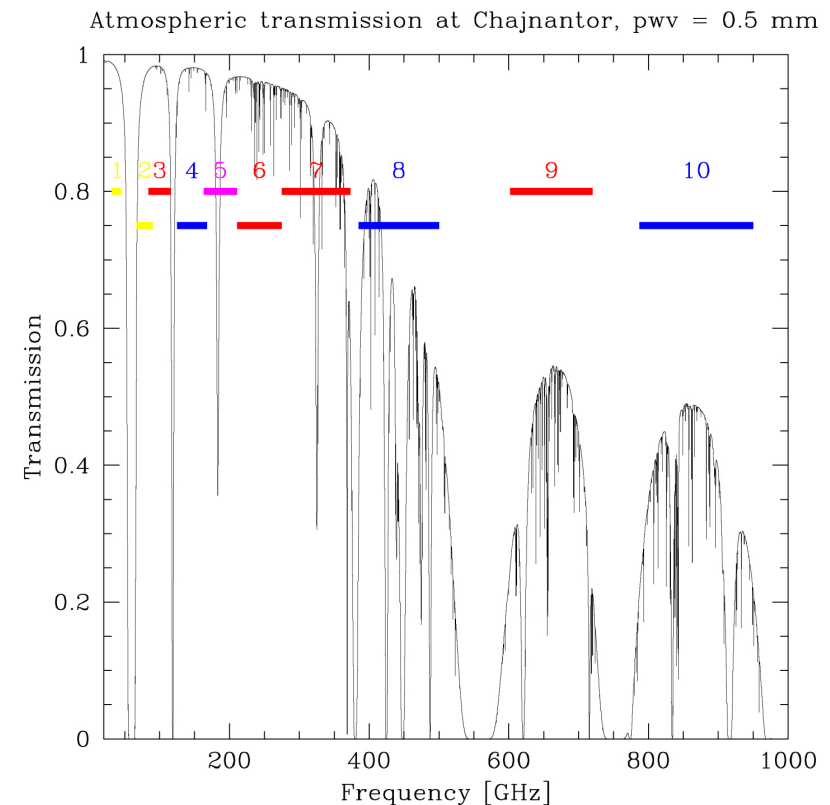
- Largest (sub)mm radio interferometer, in Chajnantor (Chile)
 - 66 antennas, at 5000 m, 0.4-3 mm, baselines up to 15 km
- Partnership between
 - ESO (37.5%)
 - NSF/AUI/NRAO (37.5%)
 - NINS/NAOJ (25%)
 - In cooperation with Chile
- Operational since 9/2011





ALMA bands & performance

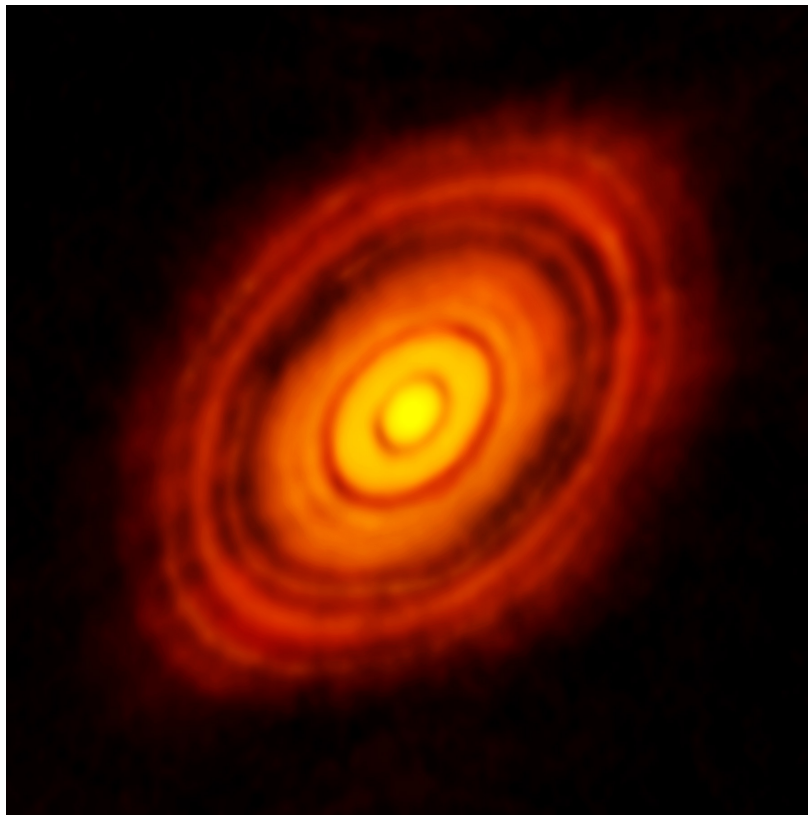
- Bands from 0.4mm to 3 mm
- Spatial resolution
 - 0.7"-4.8" in compact config
 - 6-37 mas in extended config
- Primary beam $\sim 21''$
- Spectral resolution: limited by 7680 channels and 8 GHz bandwidth.
 - 1 km/s easy to achieve
- Sensitivity: enormous



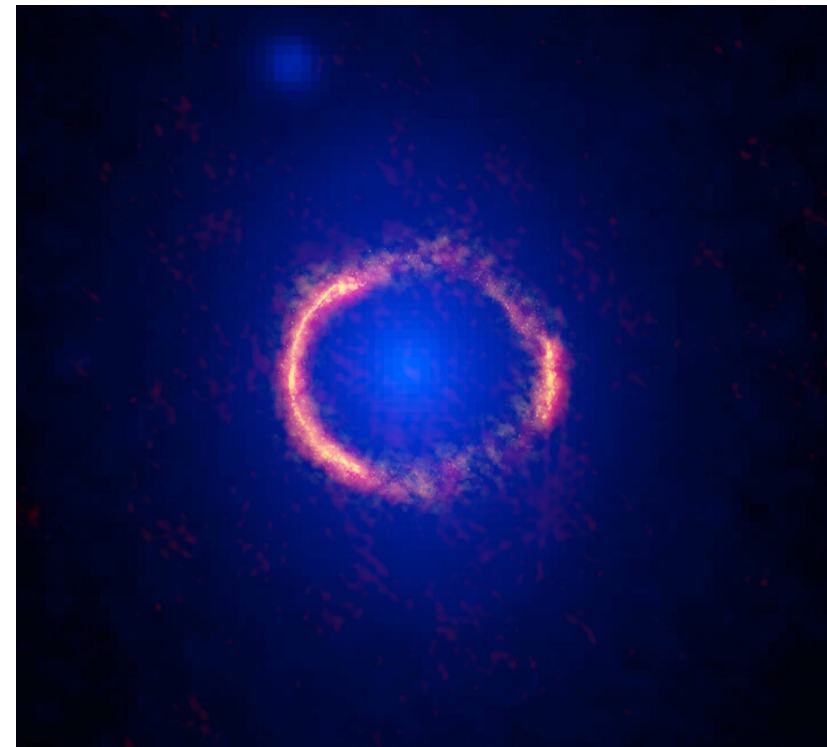
Amazing ALMA science



HL Tau proto-planetary disk



Einstein ring at $z=3$





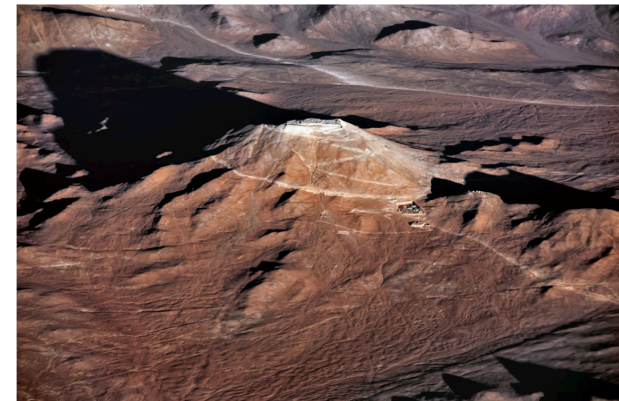
European Extremely Large Telescope (E-ELT)





The E-ELT project

- Telescope with 39.3m segmented mirror, NIR-optimized, built-in adaptive optics
- To be erected in Cerro Armazones, part of the Paranal complex
- Programme is approved, construction started in 2014. Phase 1 authorised:
 - 3 instruments funded: MICADO/MAORY, HARMONI & METIS
 - ELT-MOS & ELT-HIRES studies funded
 - Inner mirror rings not filled initially
- First light expected late 2024.

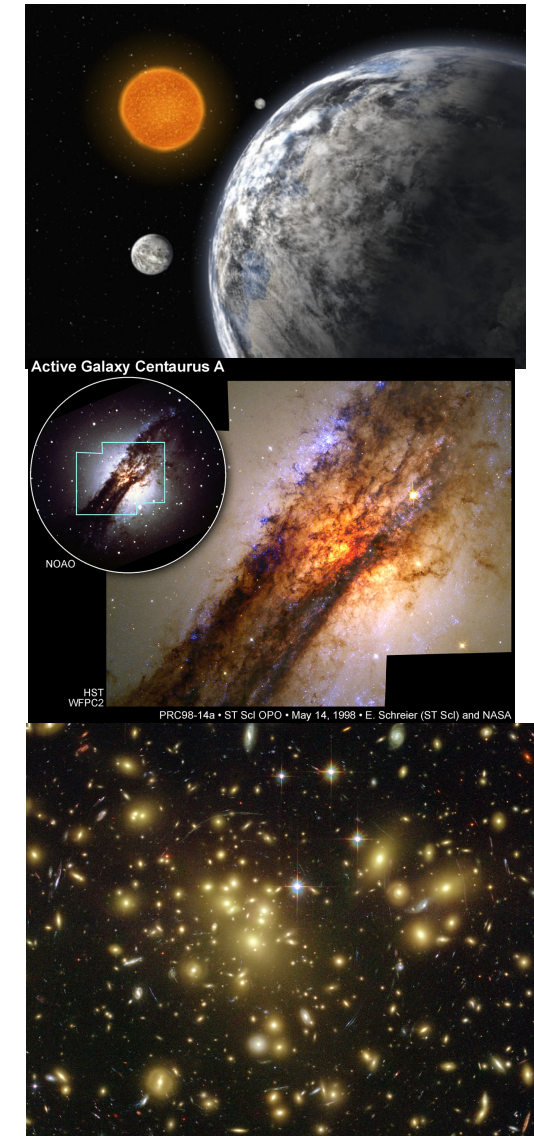




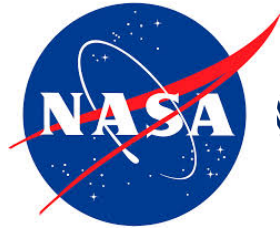
E-ELT science objectives

- Stars & planets
 - Find & characterise exo-Earths
 - Circumstellar disks
 - Stellar clusters
- Stars & galaxies
 - Stellar pops in external galaxies
 - SMBH masses in nearby galaxies
- Galaxies & Cosmology
 - First (NIR) light in the Universe
 - Chemical evolution of the Universe
 - Real-time measurement of the accelerating Universe

Exploring the Hot & Energetic Universe



ESAC, 8-10 Sep 2015

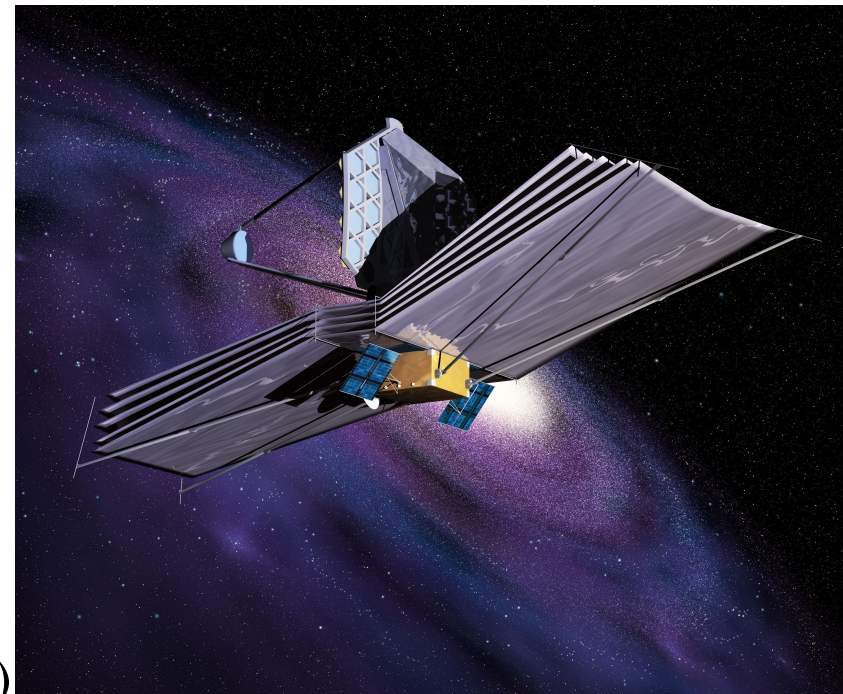


esa



JWST (2018+)

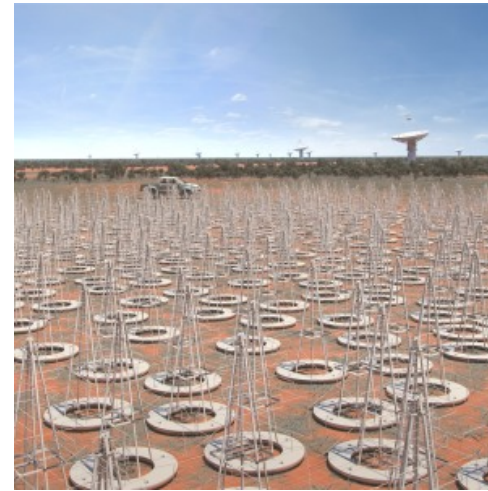
- Science objectives:
 - End of dark ages: first light & reionisation
 - Assembly of galaxies
 - Birth of stars and protoplanetary systems
 - Planetary systems & life
- Instruments
 - NIRcam (0.6-5.0 μ m, 2'x2')
 - NIRSpec (0.7-5.0 μ m, 3')
 - MIRI (5-28 μ m, 100" img, 7" spec)
 - NIRISS (0.8-2.2 μ m, 2.2')





SKA – Square Kilometre Array

- Science objectives:
 - Galaxy evolution, cosmology & dark energy
 - Strong gravity from BHs and pulsars
 - Cosmic magnetism
 - Dark ages
 - The cradle of life
- SKA1 approved (cost cap ~ 650 M€):
 - SKA1-MID (0.35-14 GHz, 200 dishes) South Africa
 - SKA1-LOW (50-350 MHz, 500 stations) Australia
 - Construction 2018-2023, early science 2020+

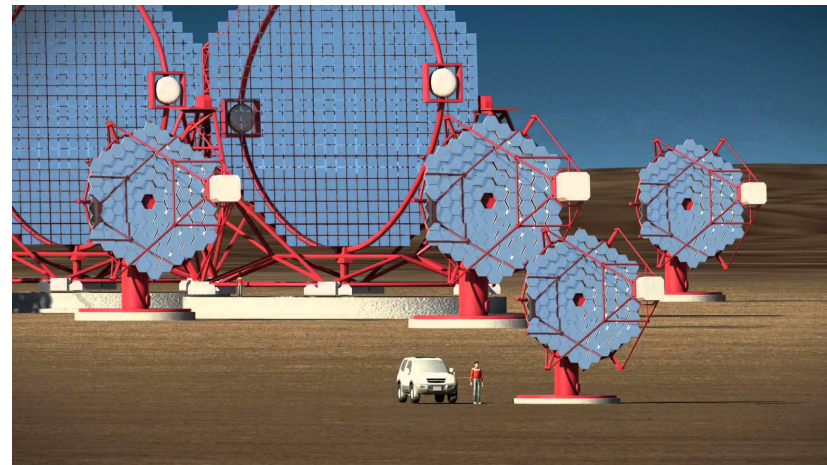
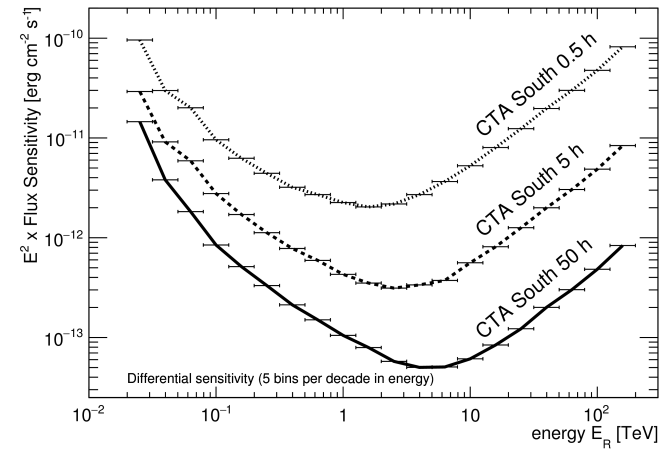




CTA – Cerenkov Telescope Array

- **First** Atmospheric Cerenkov telescope γ -ray Observatory
- Sites: La Palma (Spain) & ESO Paranal (Chile)
- Construction 2018-2023
- Dynamic range: 0.05-50 TeV
- Science objectives:
 - Origin of cosmic rays
 - Black Holes, jets and Extragalactic Background Light/SF history
 - The nature of Dark Matter & other questions in fundamental physics

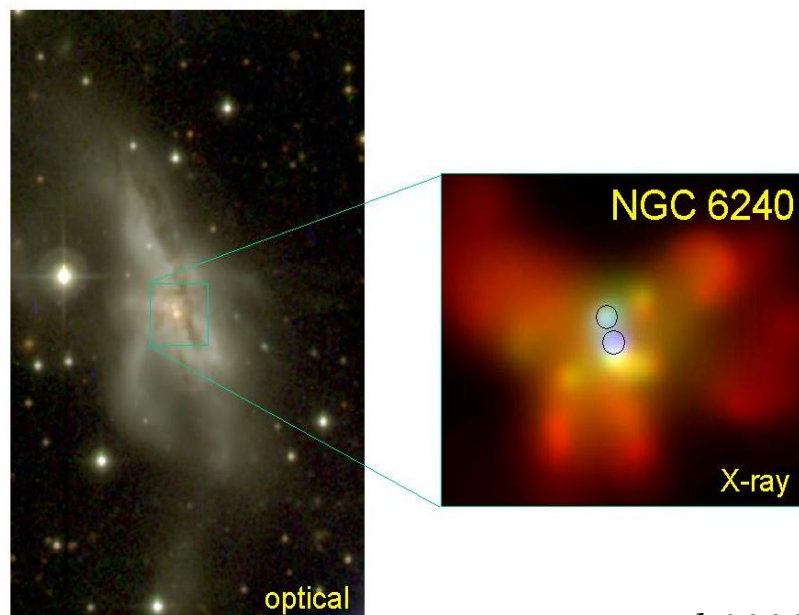
Exploring the Hot & Energetic Universe



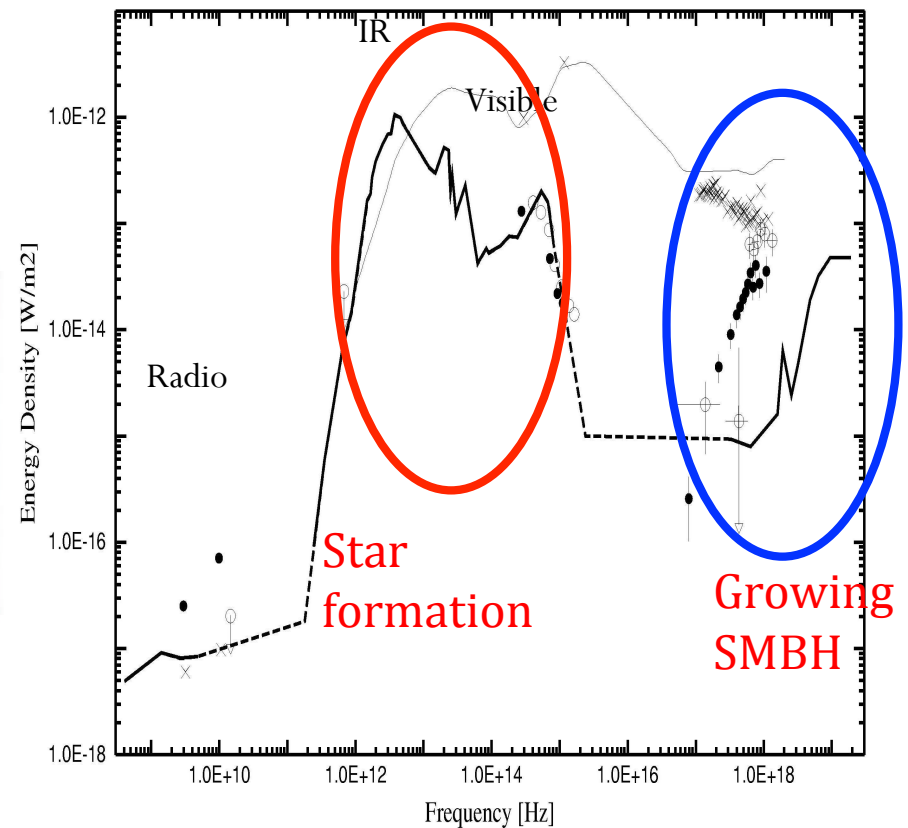
ESAC, 8-10 Sep 2015

Multi-wavelength view of galaxies

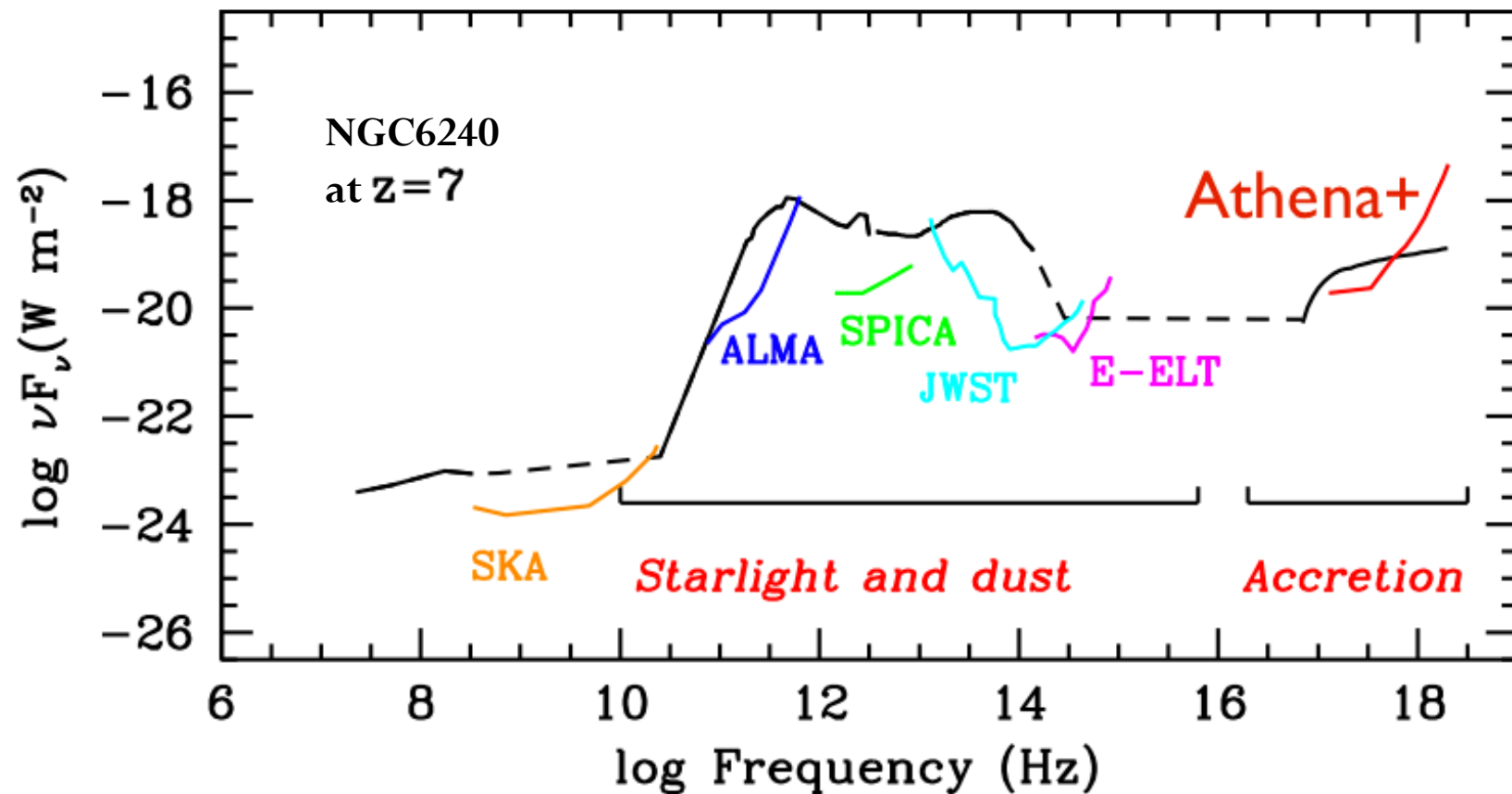
NGC6240: a SF galaxy with
a double AGN



Komossa et al 2003



Synergetic look at early galaxies

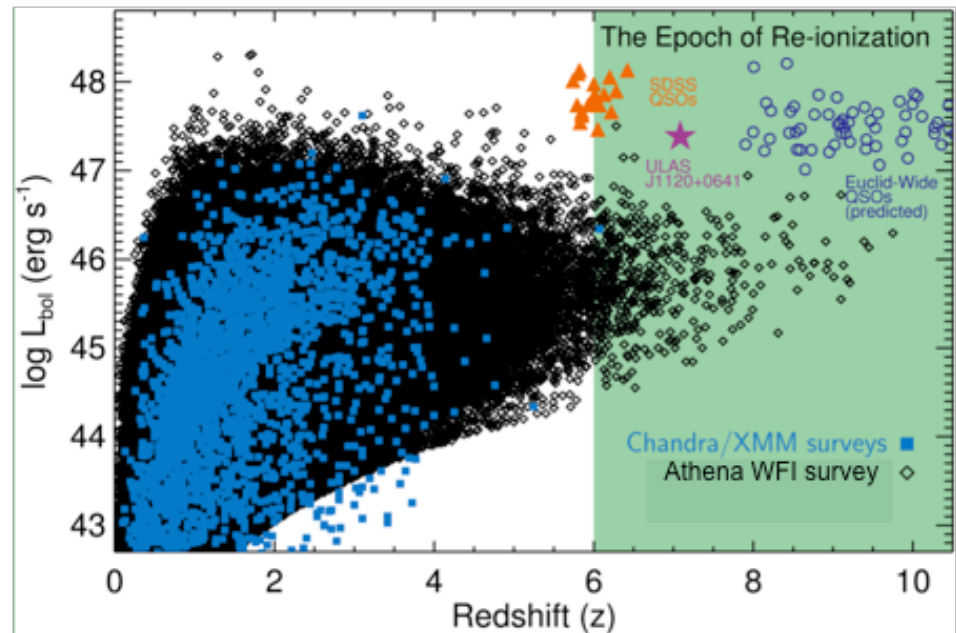


Developing synergies

- **Synergy:** Two or more agents working together to produce a result not obtainable by any of the agents independently (Source: wikipedia)
- Synergies between astronomical infrastructures are to be expected naturally: one sky, same teams
- Scientific synergies, in two flavours (and in between):
 - **Operational:** The use of other facilities is needed to achieve the own scientific goals. **Athena has several of those**
 - **Added value:** Using several facilities in conjunction leads to science added value beyond the individual goals. **Need to be identified and developed.**

High-z AGN, an example of Athena “Operational” synergy

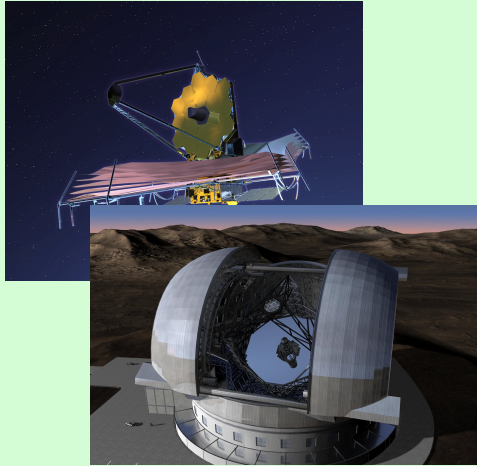
- “Needle in the haystack”:
finding $z > 6$ AGN in multi-tiered X-ray surveys:
 - Multi- λ photometry
 - Opt/NIR Spectroscopy
- This needs to be programmed carefully, and not pretend that ancillary observations will be gracefully granted.



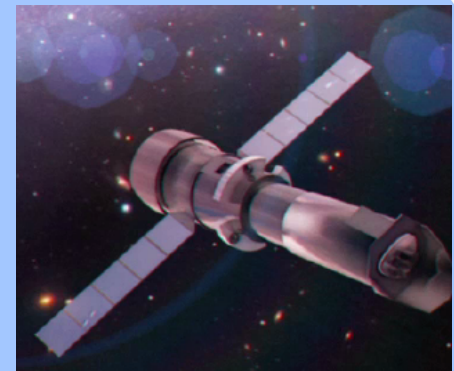
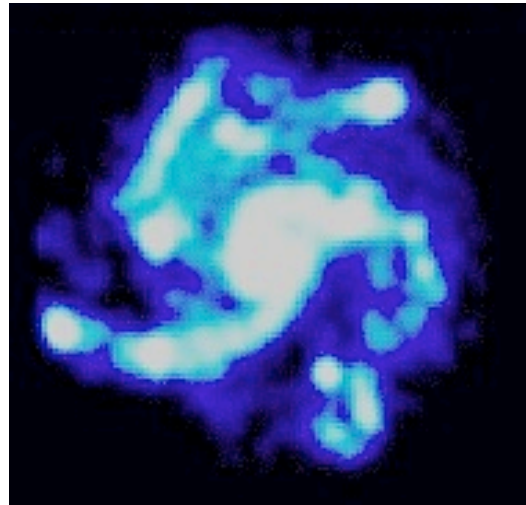
Examples of existing scientific synergies, outstanding outcome

- Cluster evolution up to $z \sim 1$ (Planck + XMM-Newton)
- Galaxy/AGN evolution (HST + Opt/NIR ground + X-ray) in selected cosmological fields.
- AGN feedback through winds & outflows (XMM-Newton + Herschel/ALMA)
- GRBs (Swift/Agile + everything)
- Molecular astrophysics (ISO/Spitzer/Herschel + ALMA + MIR instruments on ground telescopes + Rosetta)
- ...
- (Do your own ADS search and you will be surprised!)

Witnessing galaxy-SMBH co-evolution at high z



Dynamical masses of $z < 0.5$
SMBH
Stellar masses & SFR
Redshifts & morphologies



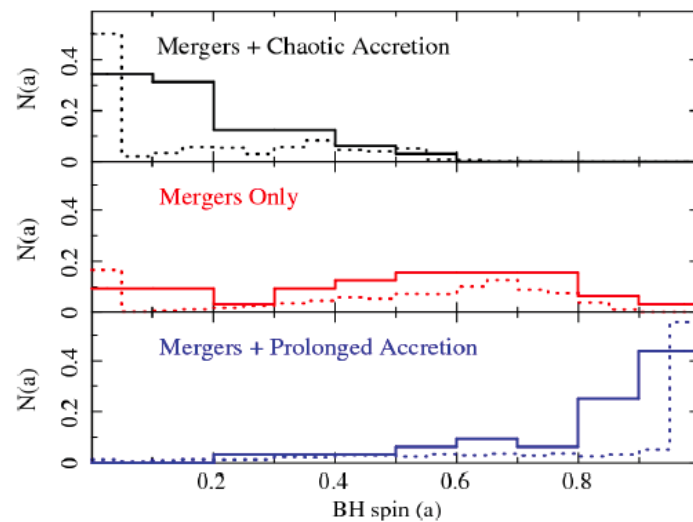
Unveil obscured AGN
Quantify AGN feedback



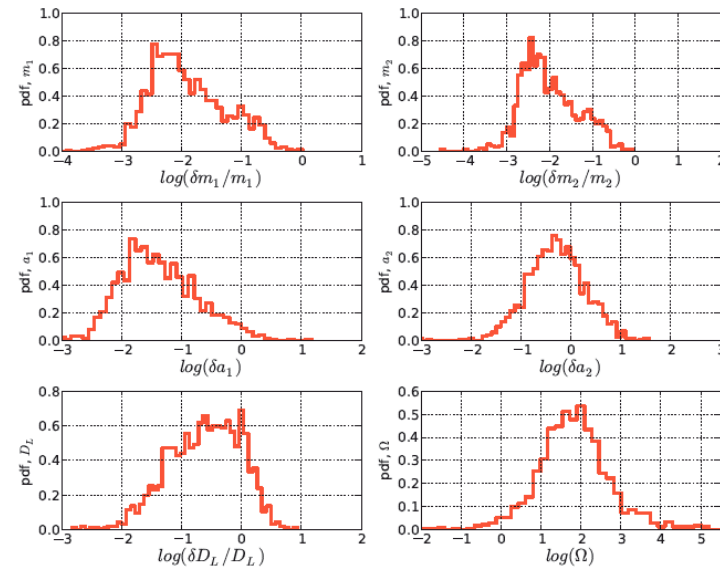
Cold gas & dust masses
Molecular outflows

SMBH spins: mergers vs accretion

X-ray Fe line profiles

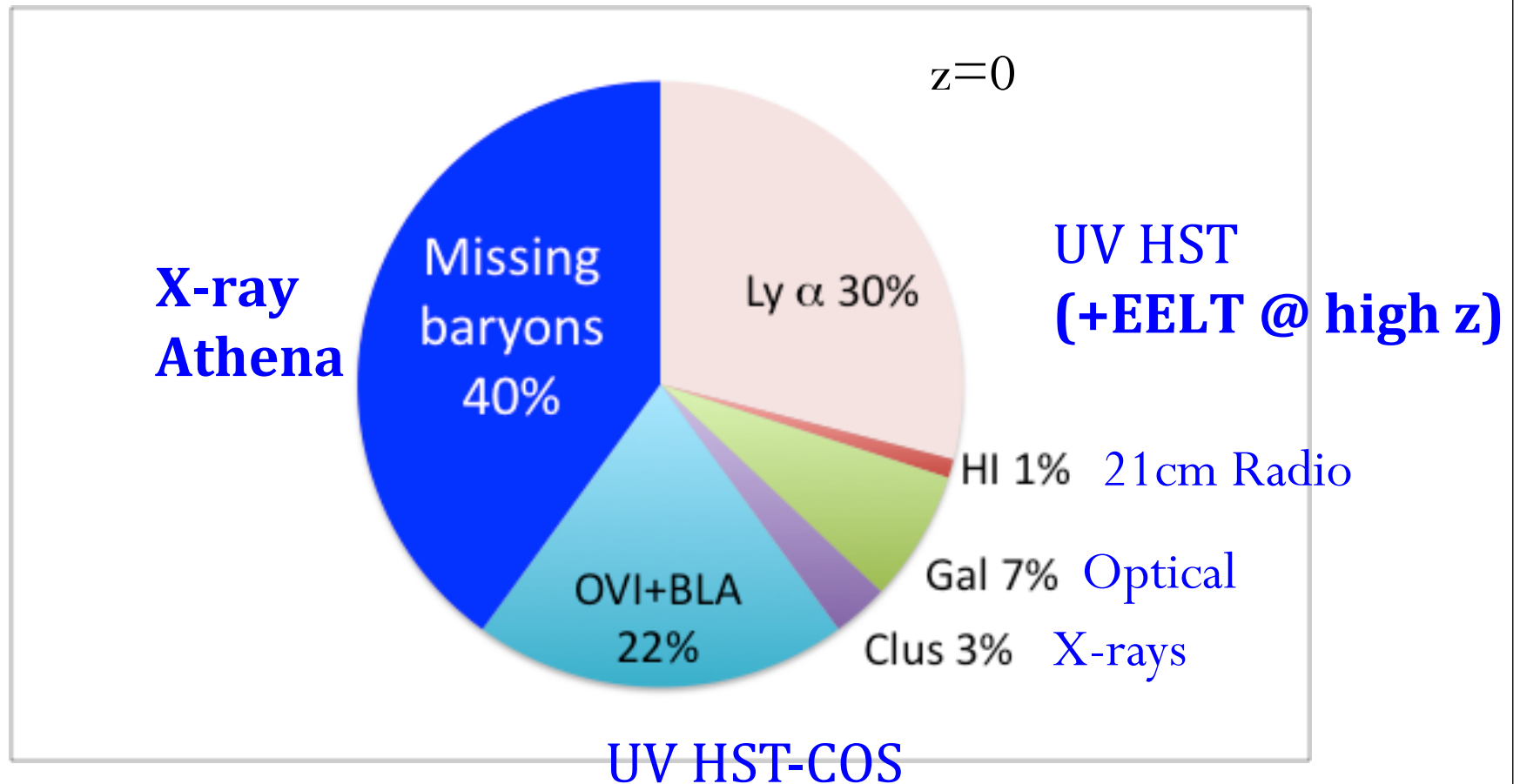


GW measurements



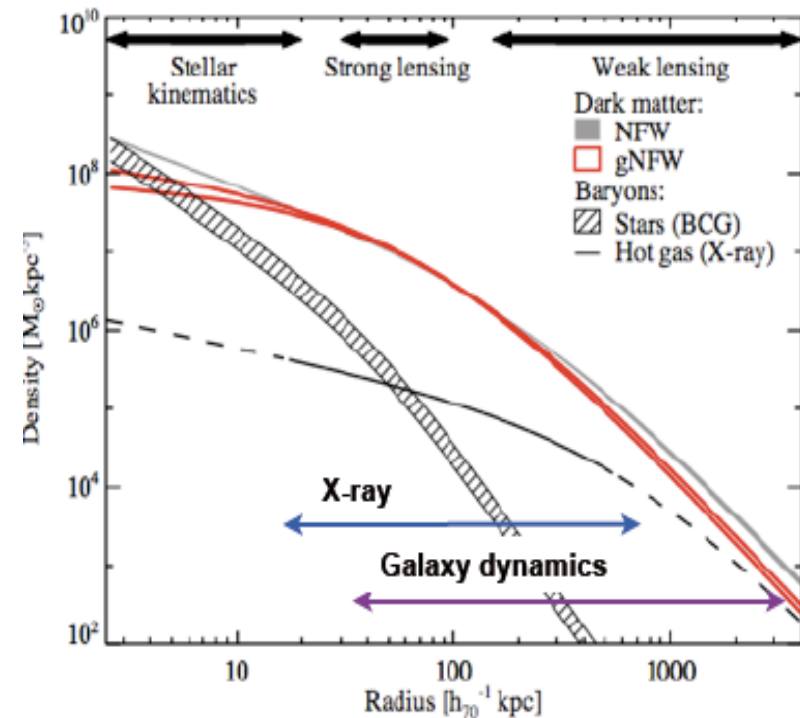
+ also from radio jets (SKA)

The cosmic baryon budget



Large Scale Structure: DM and baryons in clusters

- Mass measurements from lensing (Euclid, E-ELT) & gas hydrodynamics (Athena)
 - Measure turbulence and gas bulk motions (Athena)
- Dark Matter profile and accurate measurement of gas to mass ratio



And much more...

- Solar system planetary atmospheres (+“in situ” probes)
- Multi-messenger observations
 - CR & gamma-rays (+CTA)
 - GW signals, followed up by multi-wavelength observations, including X-rays (+VIRGO, LIGO, eLISA)
- SgrA* (+EHT, VLTI/Gravity)
- Chemical evolution (+E-ELT, ALMA)
- The transient sky (+LSST, +GRB missions)
- ...

Outlook

- Athena will be the only X-ray observatory in the late 2020s:
 - Qualitative step forward with respect to its predecessors
 - Capable of impacting all corners of Astrophysics
- Its expected performances are well matched to other contemporary facilities
 - Part of a family of powerful observatories across the spectrum
- Synergetic work of Athena with other facilities:
 - Needed to achieve science objectives (need being planned)
 - Of high added value (need being identified and developed)
- **We need to work on both types of synergies**